

Cosmology: 2015

Solutions: week 6

Exercise 1: Stress-energy tensor

a) It is convenient to split the action in its Einstein-Hilbert part and matter sector:

$$S^{EH} = \frac{1}{16\pi G} \int d^4x \sqrt{-g(x)} R$$

$$S^{matter} = \int d^4x \sqrt{-g} \left( -\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V(\phi) \right)$$

Obviously,  $S^{EH}$  is independent of  $\phi$ . Thus

$$\frac{\delta S^{EH}}{\delta \phi} = 0$$

Then

$$\frac{1}{\sqrt{-g(y)}} \frac{\delta S^{matter}}{\delta \phi(y)}$$

$$= \frac{1}{\sqrt{-g(y)}} \int d^4x \sqrt{-g(x)}$$

$$\left\{ -\frac{1}{2} g^{\mu\nu} (\partial_\mu \delta^4(x-y)) \partial_\nu \phi \right.$$

$$-\frac{1}{2} g^{\mu\nu} (\partial_\mu \phi) (\partial_\nu \delta^4(x-y))$$

$$\left. - \frac{\partial V}{\partial \phi} \delta^4(x-y) \right\}$$

The first two terms are combined, integrating by parts as always. we assume that the fields fall off at infinity, so that there are no surface- terms

$$= \frac{1}{\sqrt{-g(y)}} \int d^4x \sqrt{-g(x)} \cdot \delta^4(x-y)$$

$$\left\{ \frac{1}{\sqrt{-g(x)}} \partial_\mu (\sqrt{-g(x)} g^{\mu\nu} \partial_\nu \phi) - \frac{\partial V}{\partial \phi} \right\}$$

Carrying out the integral with the  $\delta$ -function and cancelling the determinants yields the c.o.m. for the scalar field:

$$\frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} g^{\mu\nu} \partial_\nu \phi) - \frac{\partial V}{\partial \phi} = 0 \quad (*)$$

□

For  $g_{\mu\nu} = \eta_{\mu\nu}$  we have:

$$\sqrt{-\eta} = 1 , \quad \partial_\mu \eta_{\alpha\beta} = 0$$

and

$$\frac{\partial}{\partial \phi} \left( \frac{1}{2} m^2 \phi^2 \right) = m^2 \phi$$

note: no  $\delta$ -distribution here, since by construction  $\frac{\partial V}{\partial \phi}$  is not a variation with respect to a field.

Substitute into (x):

$$\eta^{\mu\nu} \partial_\mu \partial_\nu \phi - m^2 \phi = 0$$

use  $\eta^{\mu\nu} = \text{diag}(-1, 1, 1, 1)$  to obtain:

$$(\partial_t^2 - \partial_x^2 + m^2) \phi = 0$$

- b) we first compute the variation of the Einstein - Hilbert term

$$\delta S^{EH} = \frac{1}{16\pi G} \int d^4x (\delta \sqrt{-g} R + \sqrt{-g} \delta R)$$

use identities:

$$= \frac{1}{16\pi G} \int d^4x \left\{ R \left( -\frac{1}{2} \sqrt{-g} g_{\mu\nu} \delta g^{\mu\nu} \right) \right. \\ \left. + \sqrt{-g} R_{\mu\nu} \delta g^{\mu\nu} + \dots \right)$$

$$= \frac{1}{16\pi G} \int d^4x \sqrt{-g} \left[ R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right] \delta g^{\mu\nu}$$

Thus

$$\frac{1}{\sqrt{-g}} \frac{\delta S^{EH}}{\delta g^{\mu\nu}} = \frac{1}{16\pi G} \left( R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right)$$

Per definition

$$T_{\mu\nu}^\phi = - \frac{2}{\sqrt{-g}} \frac{\delta S^{\text{matter}}}{\delta g^{\mu\nu}}$$

Thus:

$$\frac{1}{\sqrt{-g}} \frac{\delta S}{\delta g^{\mu\nu}} = 0$$

$$\begin{aligned} & \stackrel{!}{=} \frac{1}{16\pi G} (R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R) + \underbrace{\frac{1}{\sqrt{-g}} \frac{\delta S^{\text{matter}}}{\delta g^{\mu\nu}}}_{= -\frac{1}{2} T_{\mu\nu}^\phi} = 0 \end{aligned}$$

Thus

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu}^\phi$$

need to compute  $T_{\mu\nu}^\phi$  explicitly:

$$\begin{aligned} \delta g^{\mu\nu} S^{\text{matter}} &= \int d^4x \left[ (\delta \sqrt{-g}) \left( -\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V \right) \right. \\ &\quad \left. + \sqrt{-g} \left( -\frac{1}{2} (\delta g^{\mu\nu}) \partial_\mu \phi \partial_\nu \phi \right) \right] \\ &= \int d^4x \sqrt{-g} \delta g^{\mu\nu} \\ &\quad \left\{ -\frac{1}{2} g_{\mu\nu} \left( -\frac{1}{2} g^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi - V \right) - \frac{1}{2} \partial_\mu \phi \partial_\nu \phi \right\} \end{aligned}$$

Then

$$T_{\mu\nu}^\phi = \partial_\mu \phi \partial_\nu \phi - g_{\mu\nu} \left( \frac{1}{2} g^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi + V \right)$$

c) Evaluation for FRW metric

$$T_{00}^{\phi} = \dot{\phi}^2 - \frac{1}{2} \dot{\phi}^2 + V$$

$$T_{xx}^{\phi} = -a^2 (-\frac{1}{2} \dot{\phi}^2 + V)$$

we have  $u^\alpha = (1, 0, 0, 0)$  (The scalar is at rest w.r.t cosmic coordinates)

The comparison then yields:

$$S = \frac{1}{2} \dot{\phi}^2 + V$$

$$\rho = \frac{1}{2} \dot{\phi}^2 - V$$

Exercise 2: Conservation of the stress-energy tensor

a) note:

$$\begin{aligned}\delta S^{\text{matter}} &= \frac{1}{2} \int d^4x \sqrt{-g} T^{\mu\nu} \delta g_{\mu\nu} \\ &= -\frac{1}{2} \int d^4x \sqrt{-g} T_{\mu\nu} \delta g^{\mu\nu}\end{aligned}$$

use definition of  $T_{\mu\nu}$ :

$$\begin{aligned}T_{\mu\nu} &\stackrel{!}{=} -\frac{2}{\sqrt{-g}} \frac{\delta S^m}{\delta g^{\mu\nu}} \\ &= -\frac{2}{\sqrt{-g}} \left(-\frac{1}{2}\right) \int d^4x \sqrt{-g} T_{\alpha\beta} \cdot \frac{1}{2} (\delta_{\mu}^{\alpha} \delta_{\nu}^{\beta} + \delta_{\nu}^{\alpha} \delta_{\mu}^{\beta}) \\ &\quad \cdot \delta^4(x-y) \\ &= \frac{1}{\sqrt{-g}} \int d^4x \sqrt{-g} T_{\mu\nu} \delta^4(x-y) \\ &= T_{\mu\nu}(y)\end{aligned}$$

b) we start with the left-hand-side:

$$\begin{aligned}\partial_{\mu} T^{\mu\nu} &= \partial_{\mu} T^{\mu\nu} + \Gamma^{\mu}_{\mu\lambda} T^{\lambda\nu} - \Gamma^{\lambda}_{\mu\nu} T^{\mu\lambda} \\ &= \partial_{\mu} T^{\mu\nu} \\ &\quad + \frac{1}{2} g^{\mu\nu} (g_{\mu\sigma,\lambda} + g_{\lambda\sigma,\mu} - g_{\mu\lambda,\sigma}) T^{\lambda\nu} \\ &\quad - \frac{1}{2} g^{\lambda\sigma} (g_{\mu\sigma,\nu} + g_{\nu\sigma,\mu} - g_{\mu\nu,\sigma}) T^{\mu\lambda} \\ &= \partial_{\mu} T^{\mu\nu} + \frac{1}{2} g^{\mu\nu} (g_{\mu\sigma,\lambda}) T^{\lambda\nu} - \frac{1}{2} g_{\mu\sigma,\nu} T^{\mu\sigma}\end{aligned}$$

compare to right-hand-side

$$\frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} T^{\mu\nu}) - \frac{1}{2} T^{\alpha\beta} (g_{\alpha\beta,\nu})$$

$$= \partial_\mu T^{\mu\nu} + \frac{1}{2} g^{\alpha\beta} (g_{\alpha\beta,\mu}) T^{\mu\nu} - \frac{1}{2} T^{\alpha\beta} (g_{\alpha\beta,\nu})$$

= LHS

note that we used a variant of eq. (3) replacing  $\delta \rightarrow \partial_\mu$

$$\partial_\mu \sqrt{-g} \stackrel{3b}{=} \frac{1}{2} \sqrt{-g} g^{\alpha\beta} (\partial_\mu g_{\alpha\beta})$$

c) start from eq. (9) and substitute (12):

$$\delta S^{\text{matter}} = \frac{1}{2} \int d^4x \sqrt{-g} T^{\mu\nu} \delta g_{\mu\nu}$$

$$= \frac{1}{2} \int d^4x \sqrt{-g} T^{\mu\nu} (g^\alpha g_{\mu\nu,\alpha} + (\partial_\mu g^\alpha) g_{\alpha\nu} + (\partial_\nu g^\alpha) g_{\mu\alpha})$$

$$= \frac{1}{2} \int d^4x \sqrt{-g} T^{\mu\nu} (g^\alpha g_{\mu\nu,\alpha} + 2 g_{\alpha\nu} (\partial_\nu g^\alpha))$$

$$\text{P.I.} = \int d^4x \sqrt{-g} g^\alpha \left\{ \frac{1}{2} T^{\mu\nu} g_{\mu\nu,\alpha} \right.$$

$$\left. - \frac{1}{\sqrt{-g}} \partial_\nu (\sqrt{-g} T^{\nu\alpha}) \right\}$$

$$= - \int d^4x \sqrt{-g} g^\alpha \left\{ \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} T^{\mu\nu}) - \frac{1}{2} T^{\mu\nu} g_{\mu\nu,\alpha} \right\}$$

$$= - \int d^4x \sqrt{-g} g^\alpha (\partial_\mu T^{\mu\nu})$$

d)

If  $\delta_c$  corresponds to a coordinate transformation and  $S^{\text{matter}}$  is invariant, we have  $\delta_c S^{\text{matter}} = 0$

Since  $f^\alpha$  is an arbitrary function, this can only be true if

$$D_\mu T^{\mu\alpha} = 0$$

$\Rightarrow$  conservation law for the energy momentum tensor based on symmetry principles.